

LAMOST Spectroscopic Survey

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Present by
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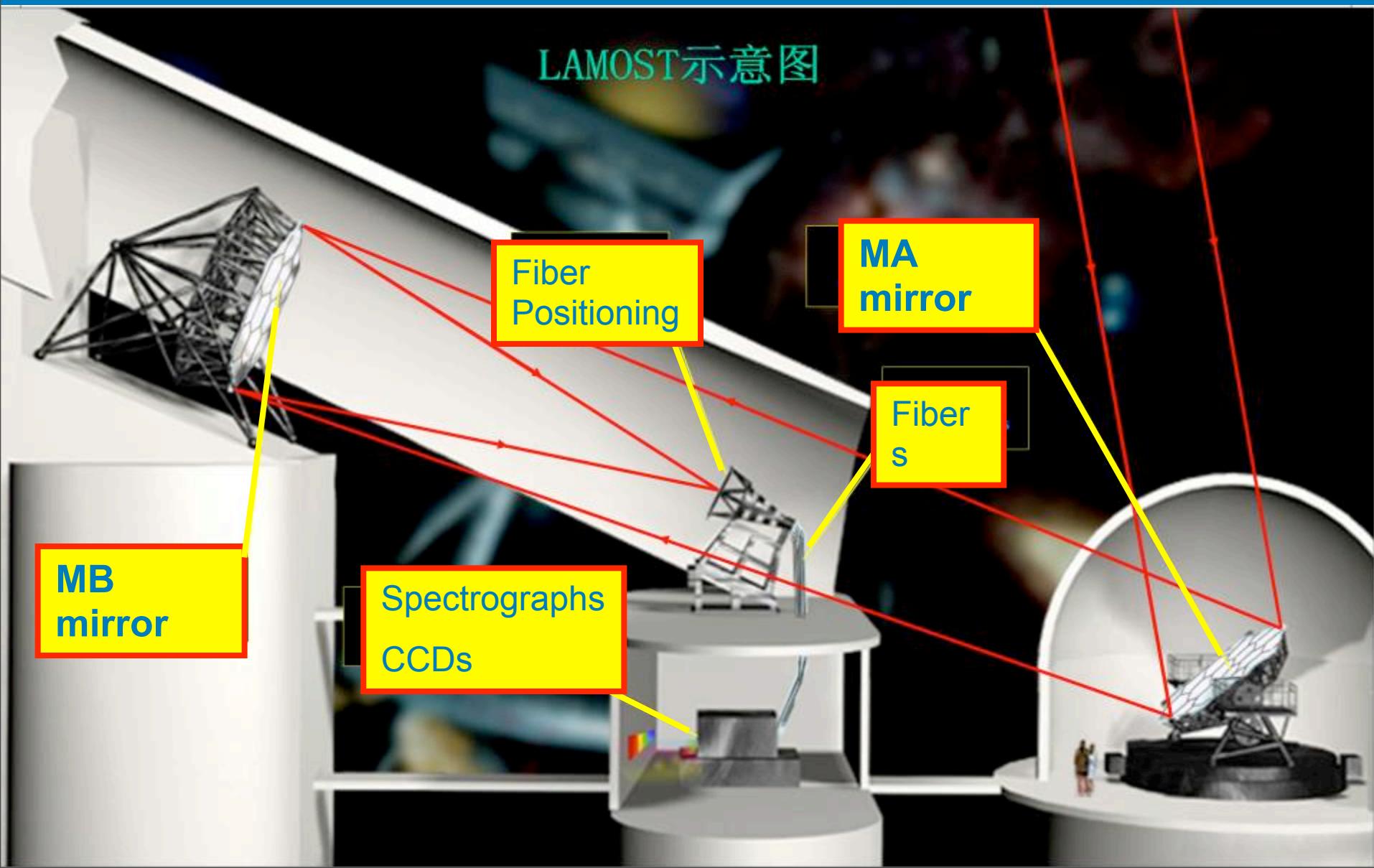
Large Sky Area Multi-Object Fiber Spectroscopic Telescope (LAMOST)

- A meridian active reflecting Schmidt telescope
- Started in 1997
- First light in August 2008
- Inauguration in 16 October 2008
- now in commissioning stage

Characteristics of LAMOST

- Effective aperture **4 meter**
- FOV **5° (1.75m linear)**
- Number of optical fiber **4000**
- Observing sky area **-10° ≤ δ ≤ +90°**
- Spectral resolution **1-0.25nm**
- Size of fiber **3.30 arcsec (320 macro linear)**
 - Site seeing: ~2 arcsec
- Survey capability **taking spectral resolution 1nm,
integration time 1.5 hours,
magnitude limit: 20.5^m**

Structure of LAMOST

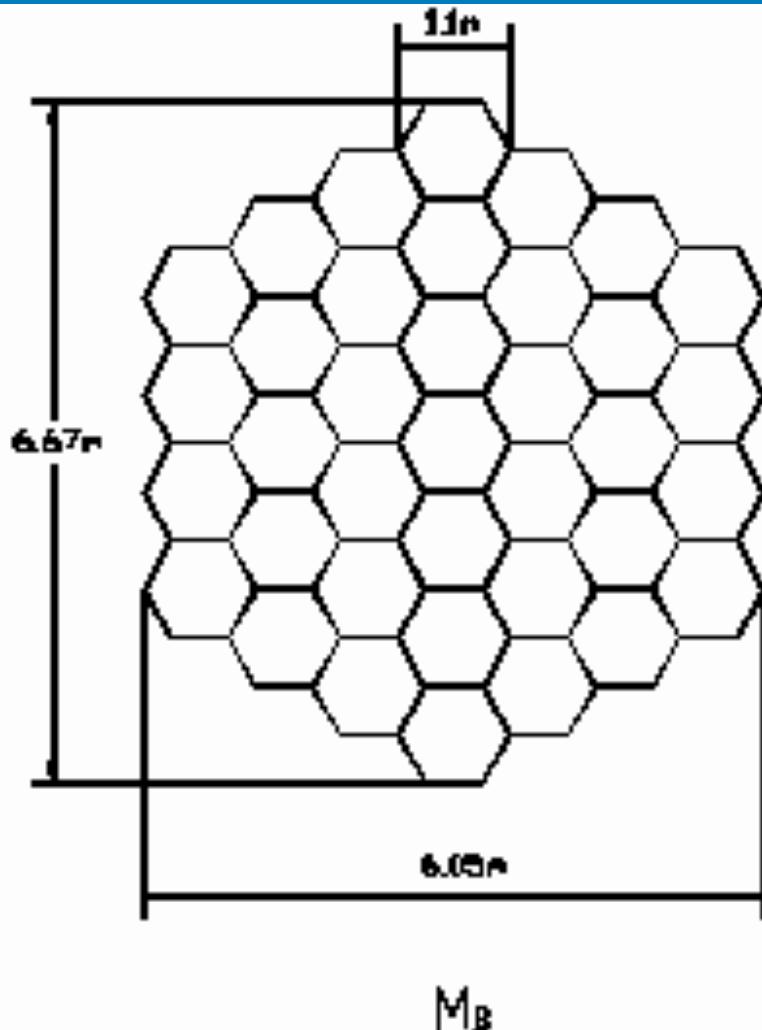
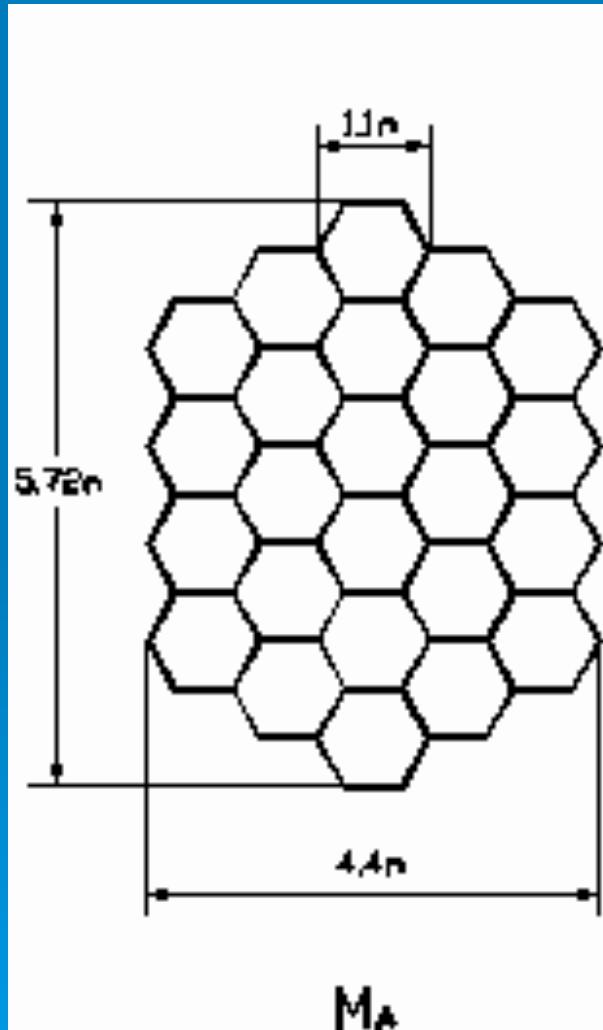


Enclosure

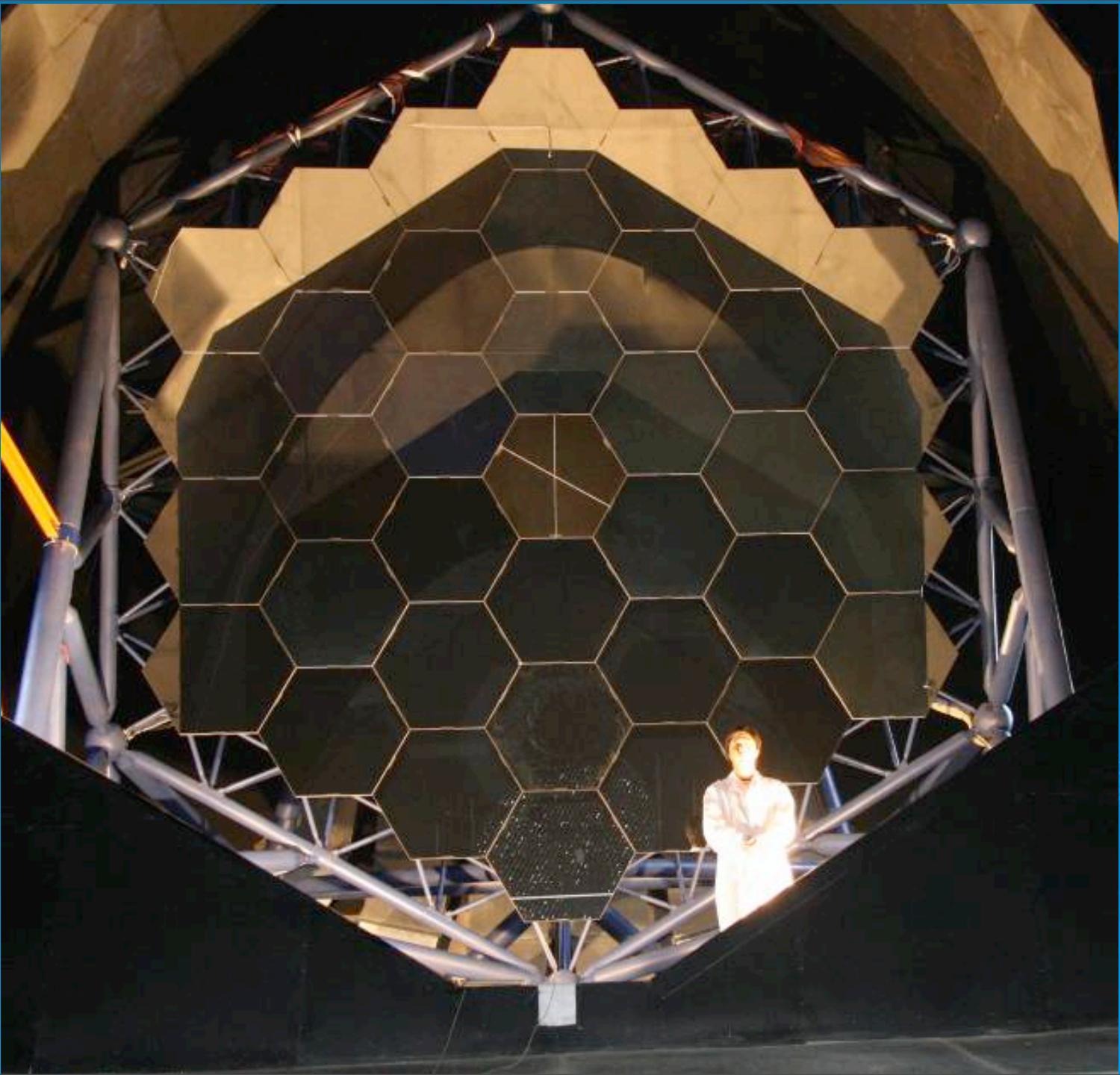


Oct 20, 2007

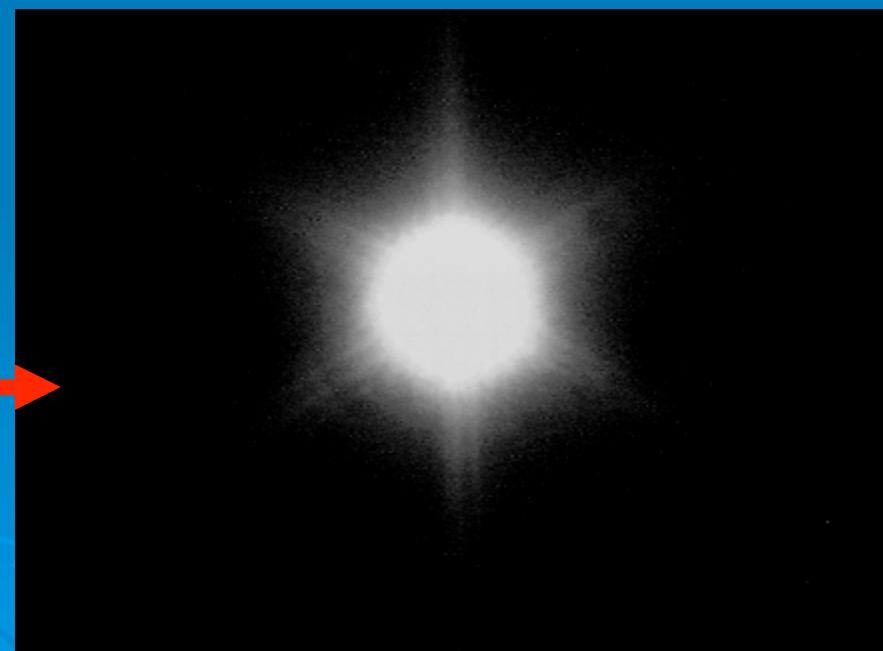
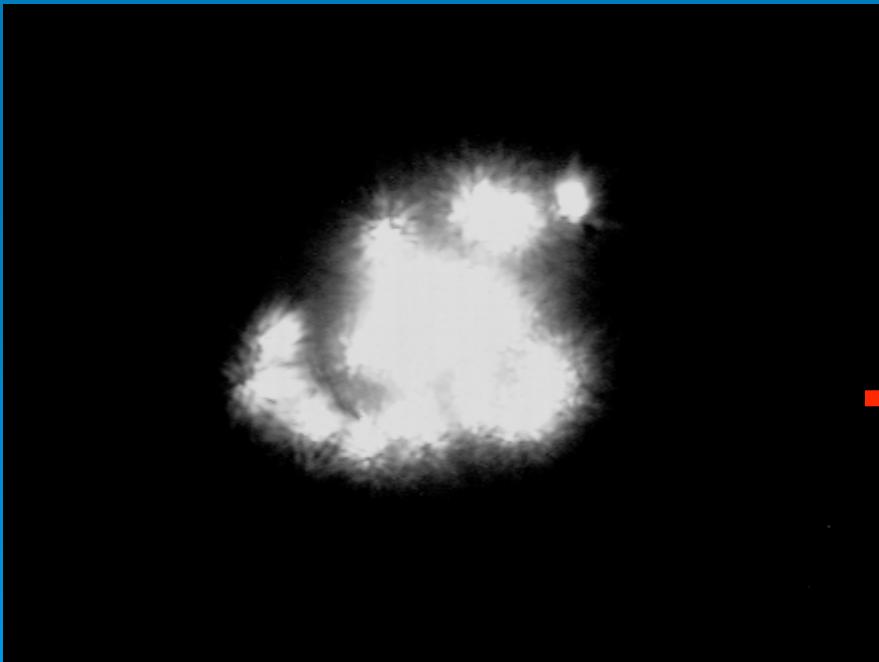
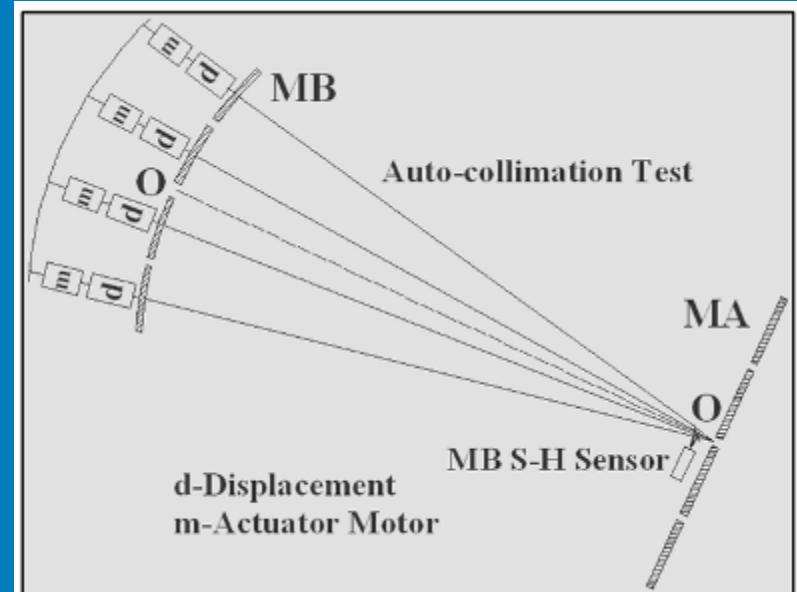
- MA: reflecting corrector (24 sub-mirrors) $\sim 4.9\text{m}$
- MB: spherical mirror (37 sub-mirrors) $\sim 6.1\text{m}$



MB

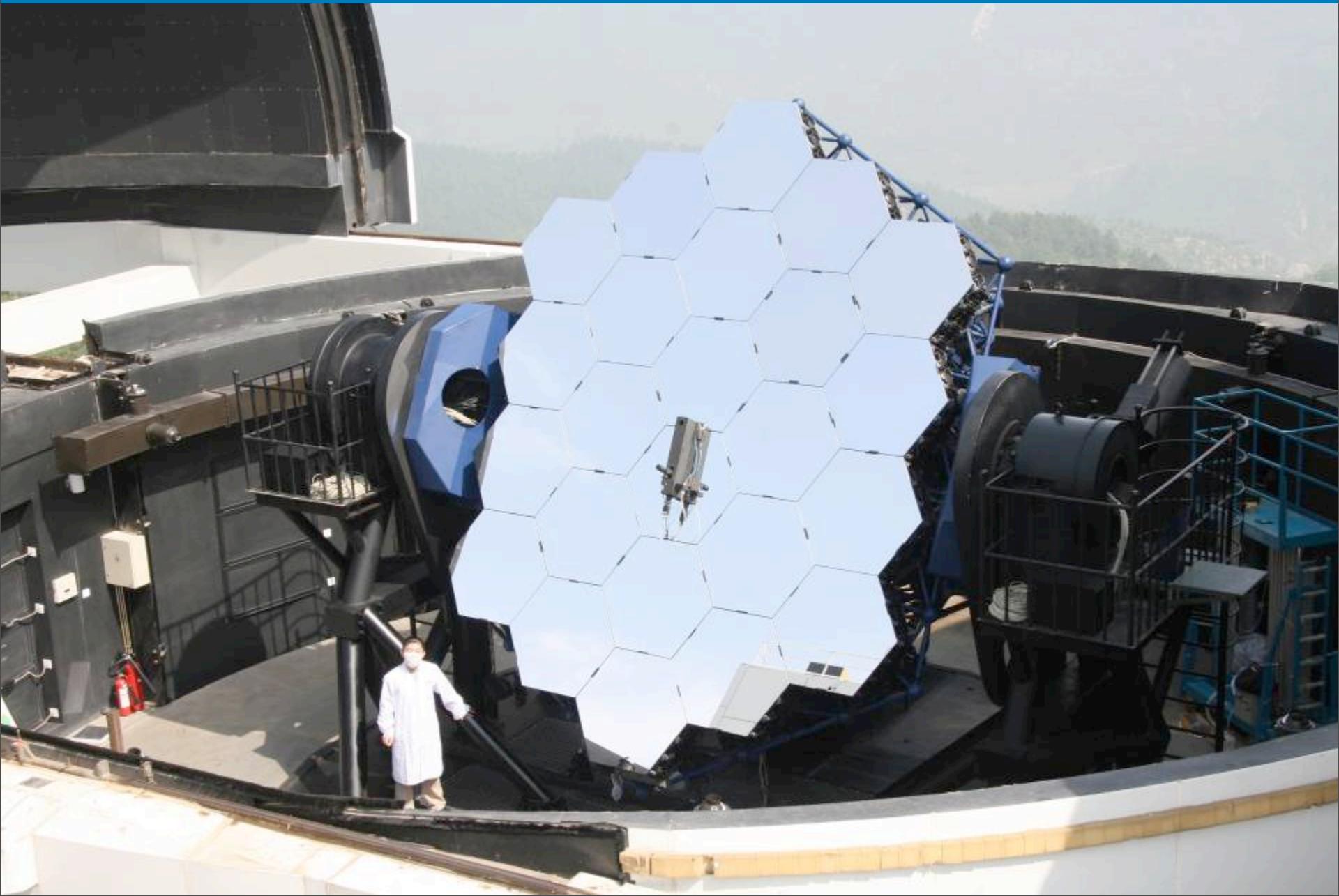


➤ Segmented Active optics for 37 sub-mirrors of MB

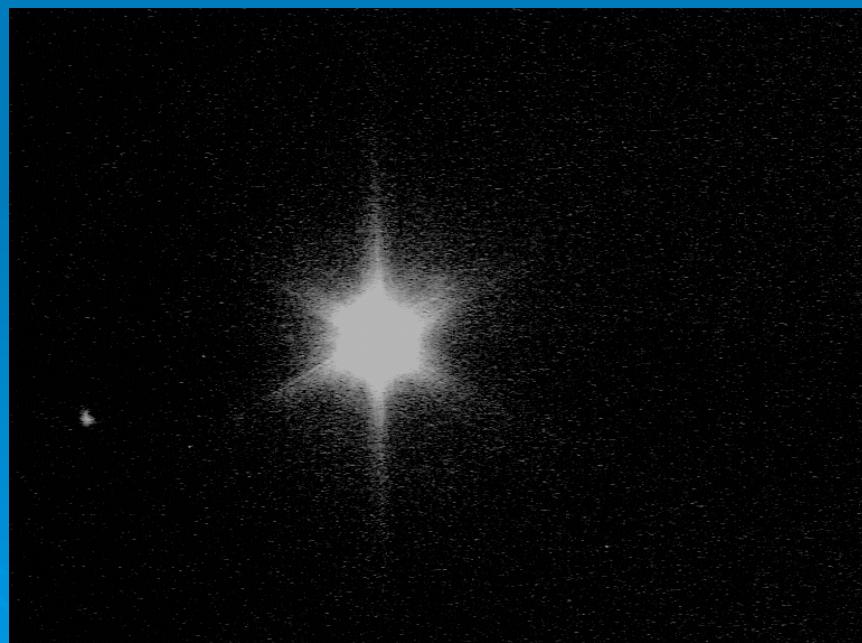
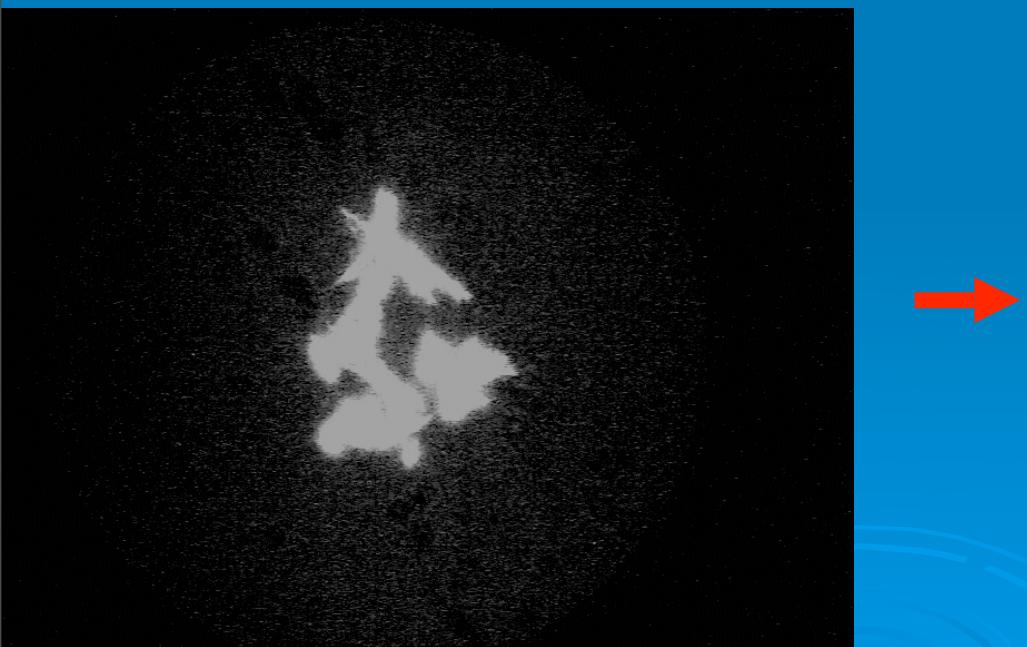
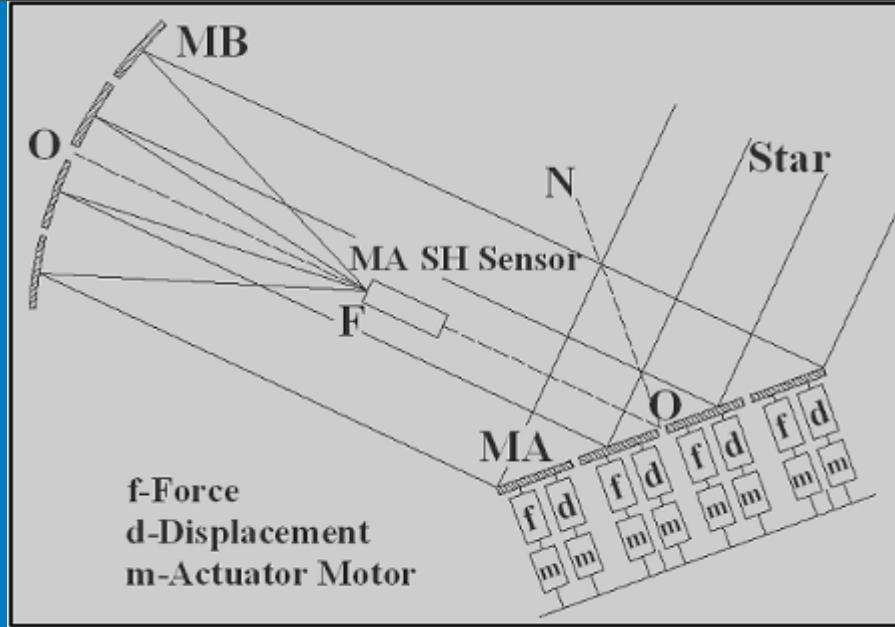


(July 13,2008)

24 sub-mirrors of MA

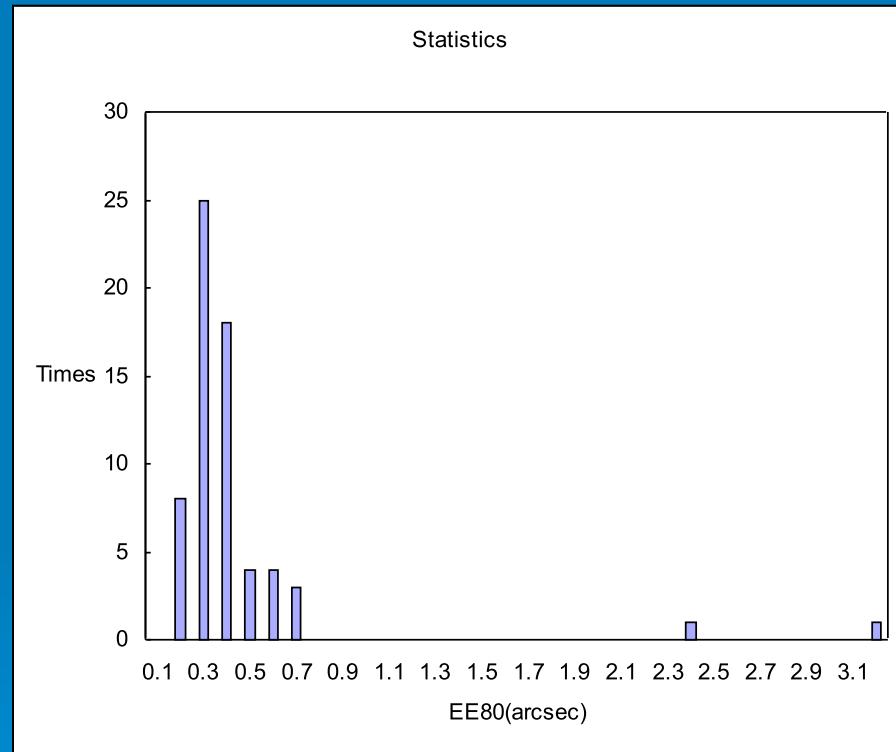
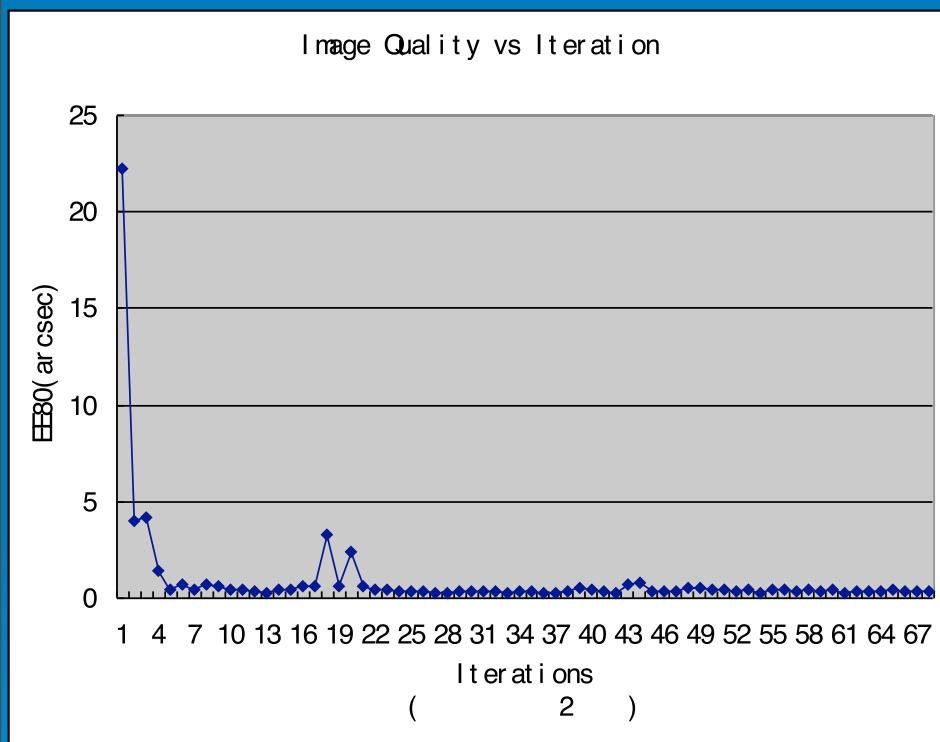


- Segmented and thin mirror active optics for 24 sub-mirrors of MA



(Sept. 10, 2008)

Image quality of LAMOST



(Nov. 21, 2008)

Instruments



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- 4000 Fibers (130km)

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- 4000 Fiber positioning units
 - 8000 step motors
- 16 Spectrographs
 - 250 fibers per spectrograph

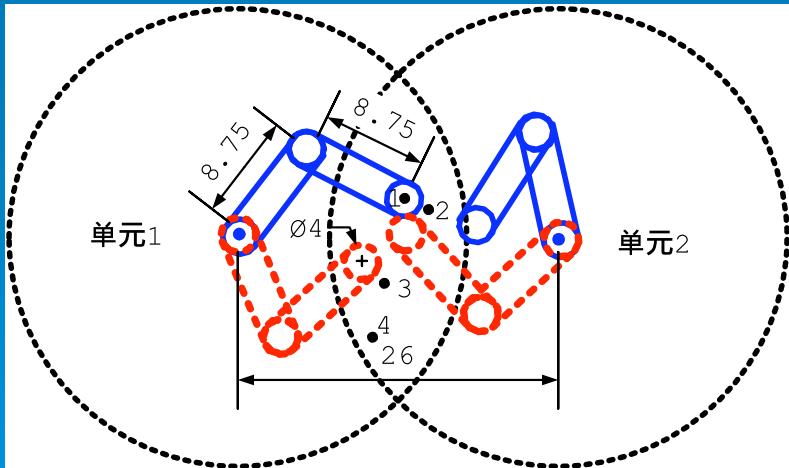
Instruments

- 4000 Fibers (130km)
- 4000 Fiber positioning units
 - 8000 step motors
- 16 Spectrographs
 - 250 fibers per spectrograph
- 32 4k x 4k CCD Cameras
 - E2V CCD chips

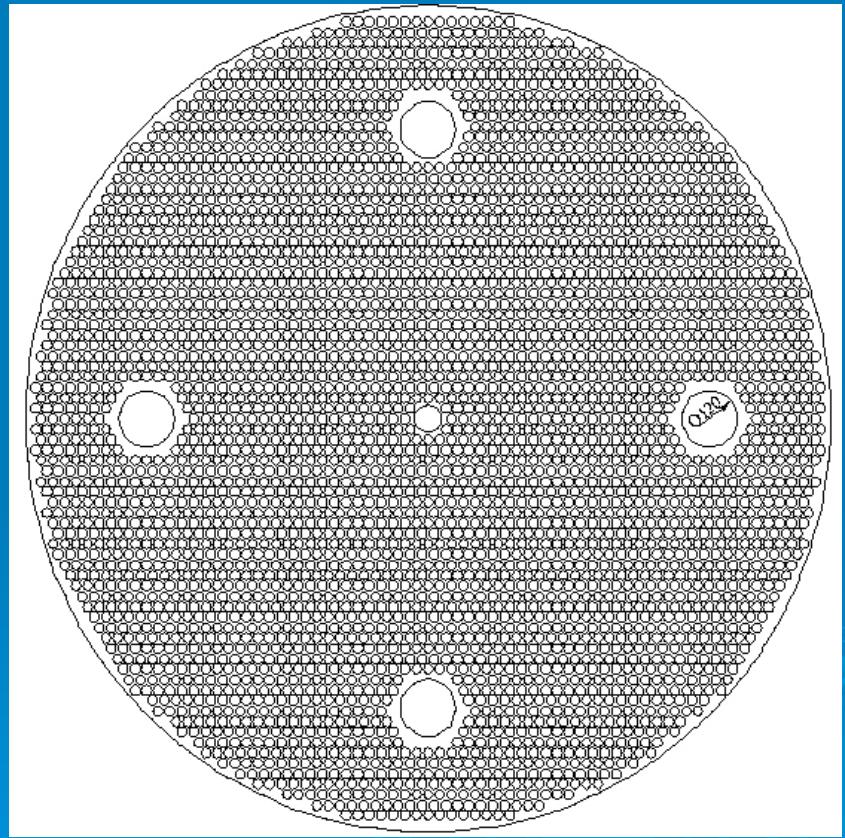
4000 fiber positioning units



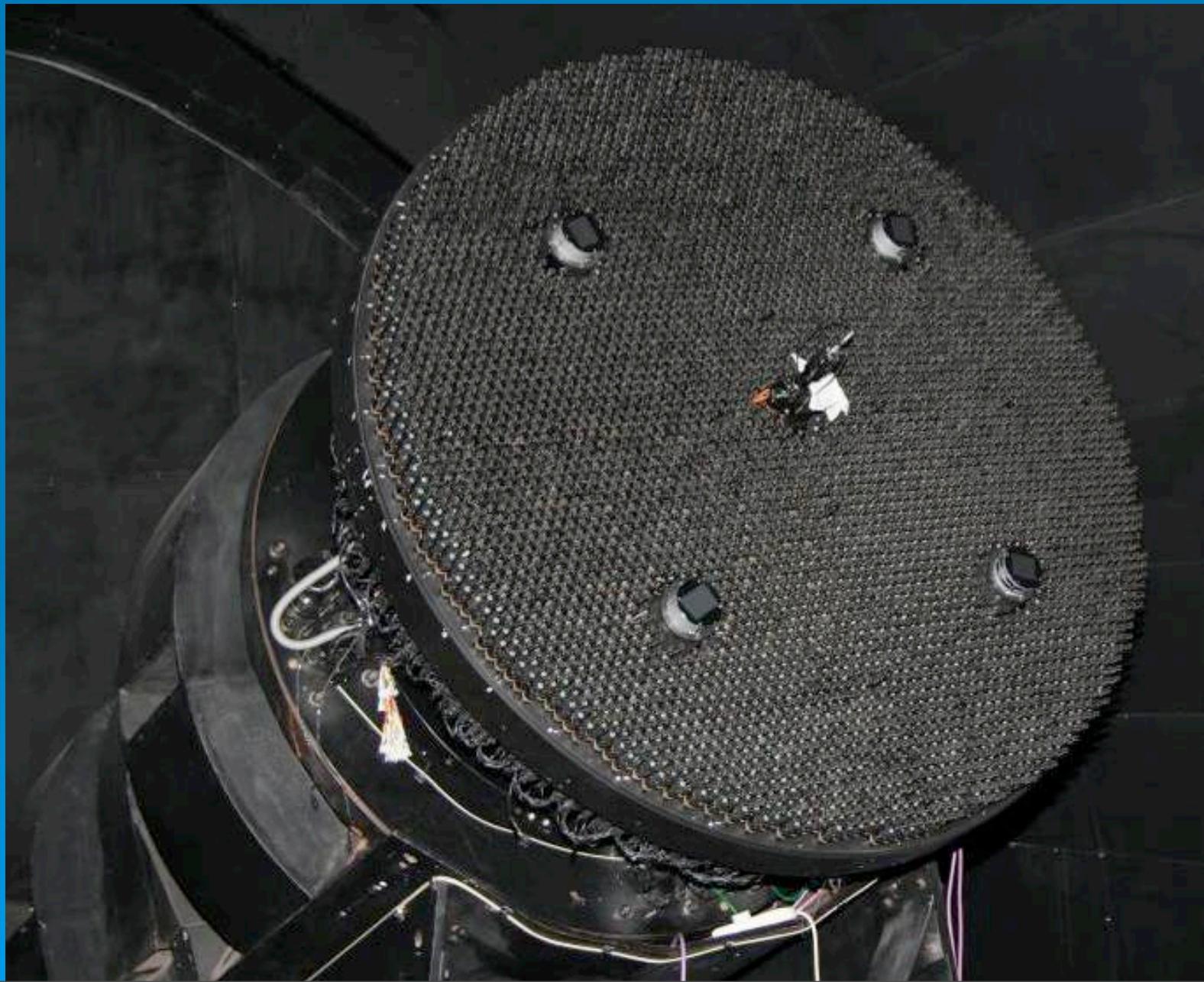
Positioning unit with 2 step motors

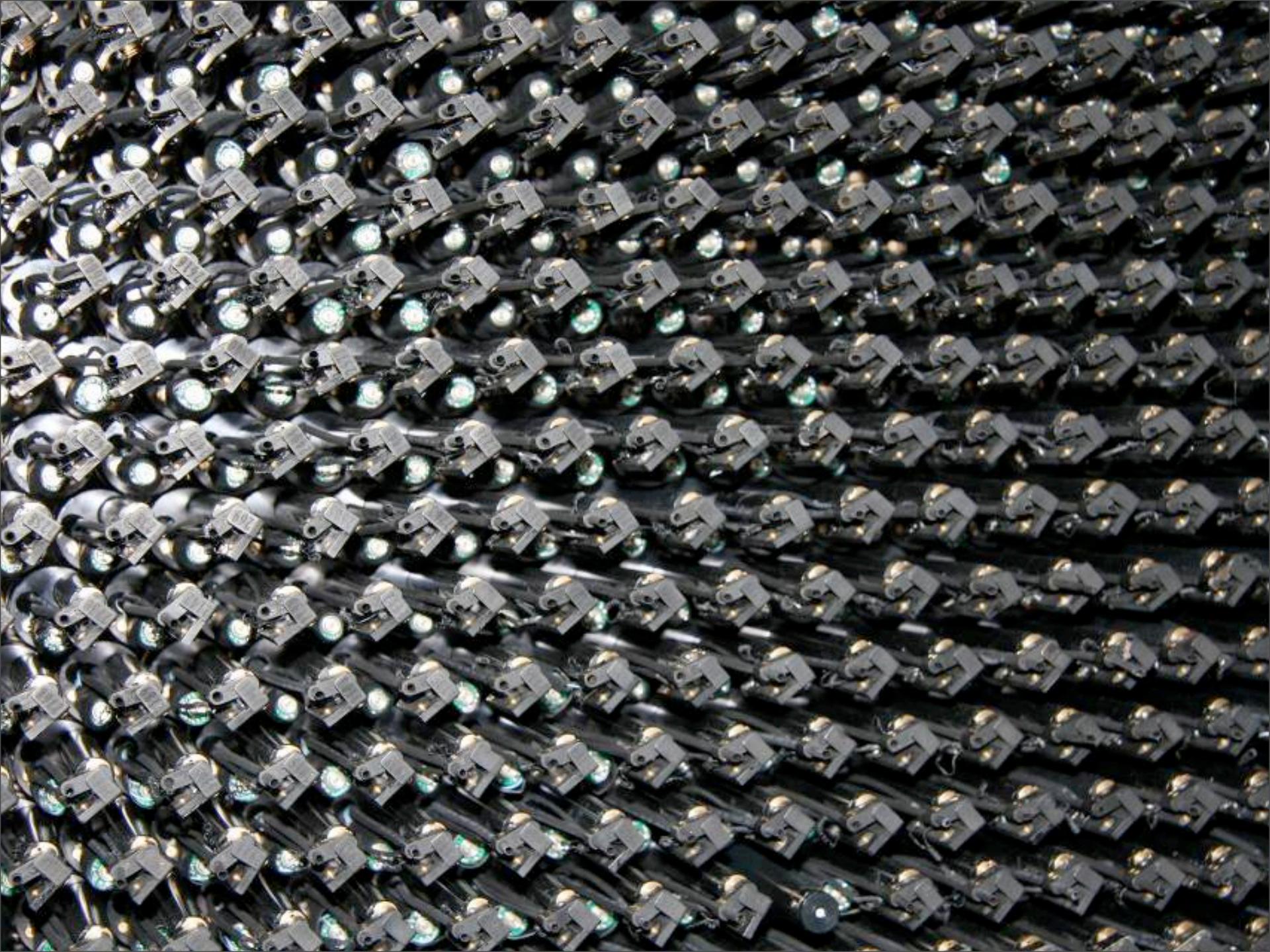


Double arm scheme



Focal Plate for holding 4000 fibers



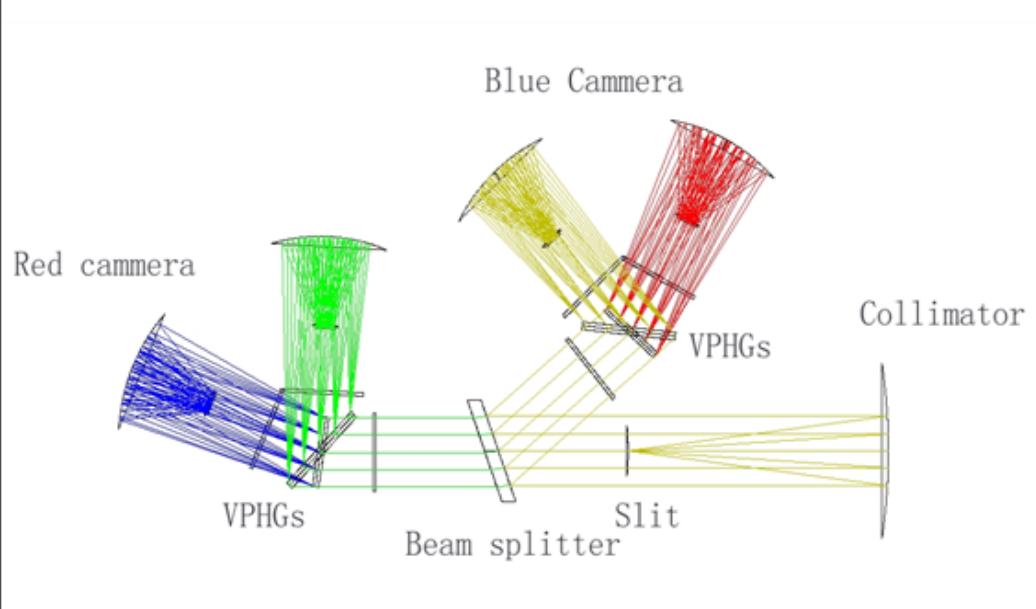
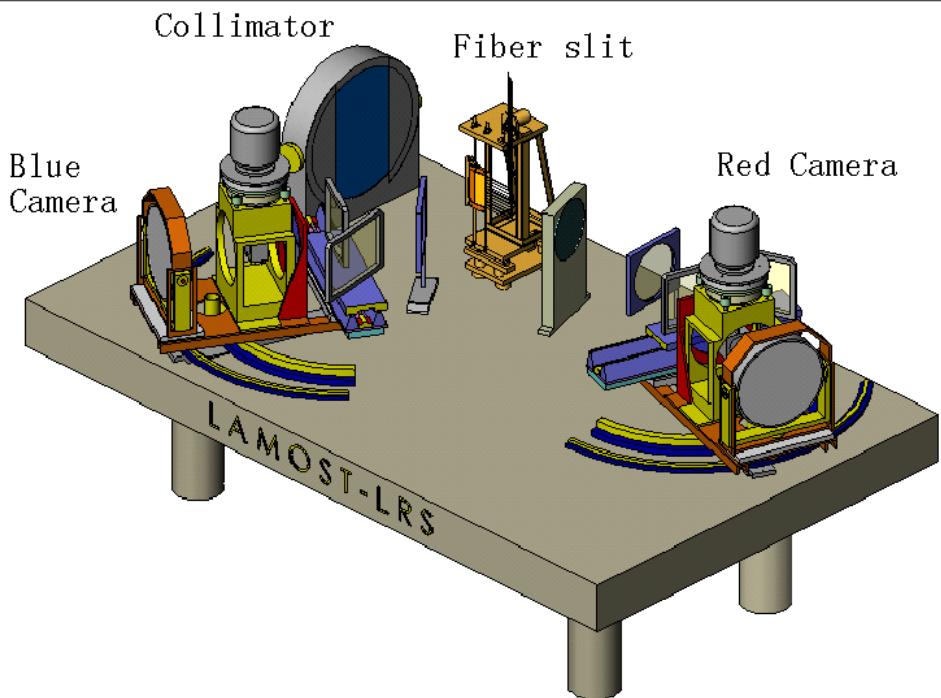


16 spectrographs

250 fibers per spectrograph

$R_L = 1000/2000$

$R_M = 5000/10000$



Spectral range:

Low blue: 370—590nm

red: 570—900nm

Medium blue: 510nm — 540nm

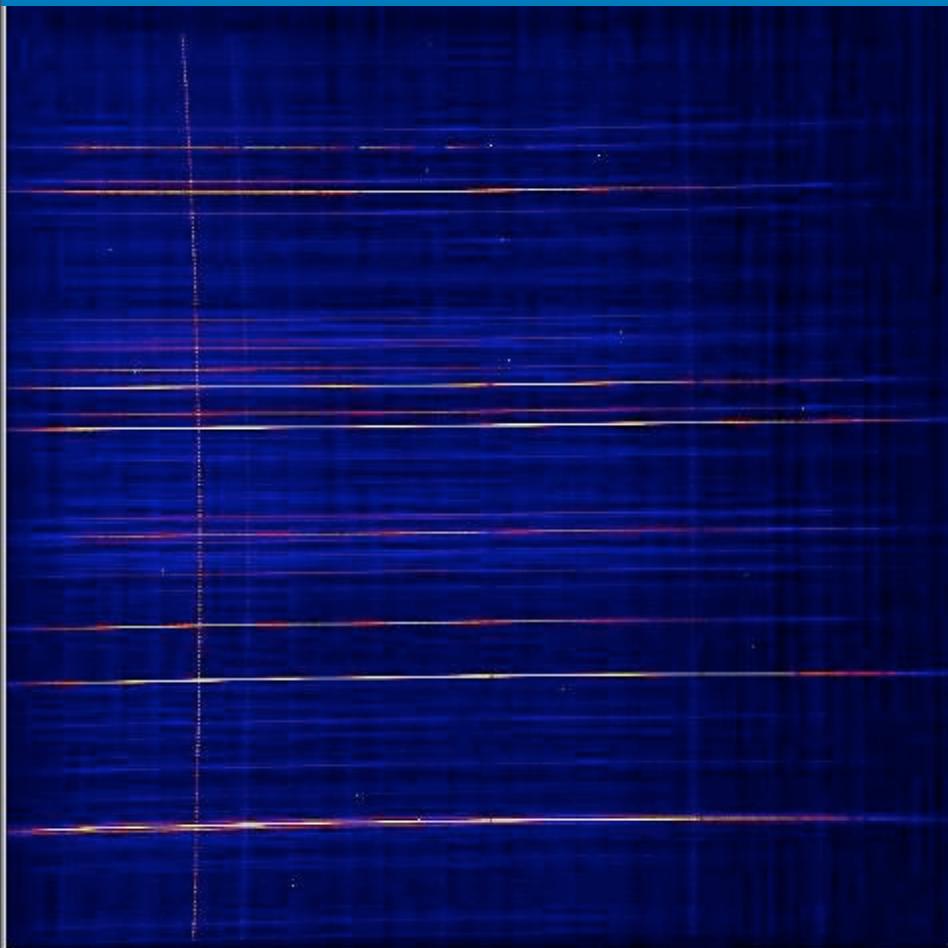
red: 830nm — 890nm



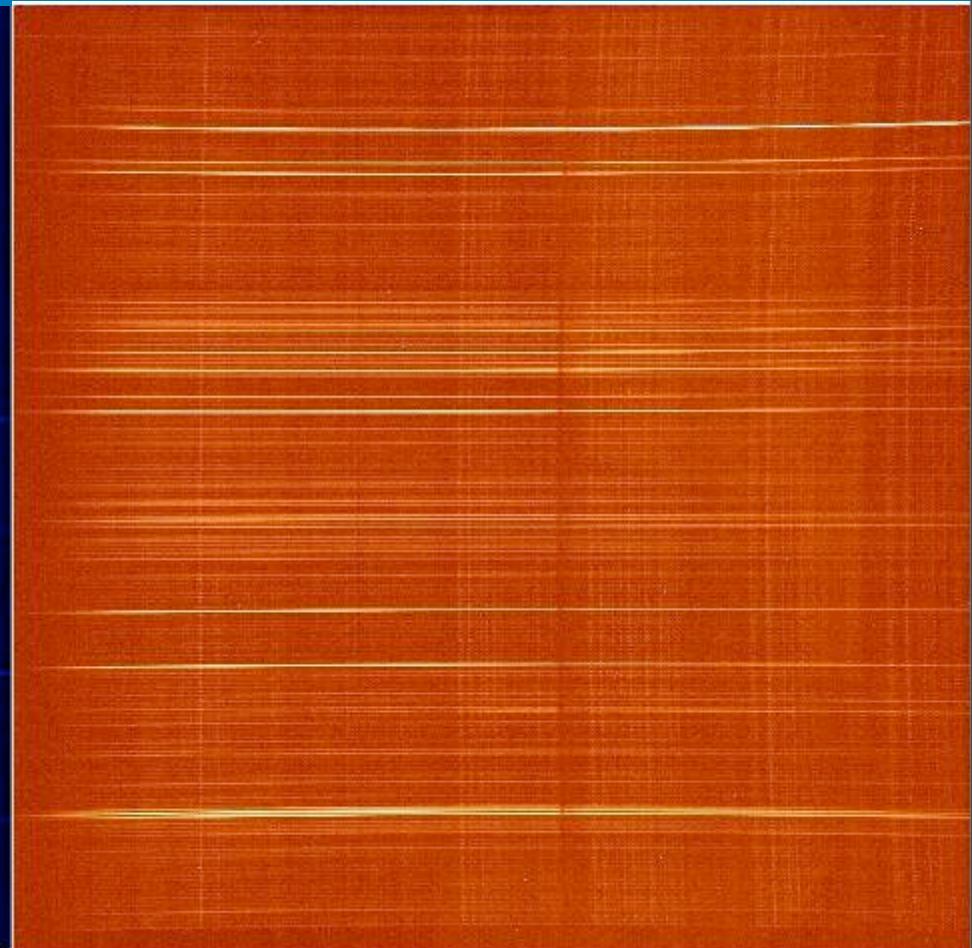
Test Observations

- Sept. 28, 2008
 - More than 2000 spectra of bright stars got in one test observation
- Dec. 27, 2008
 - M31

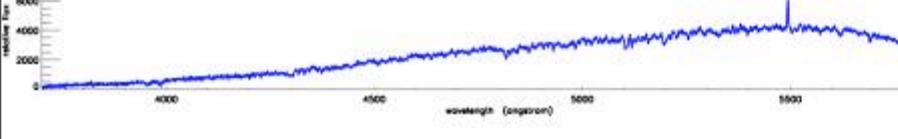
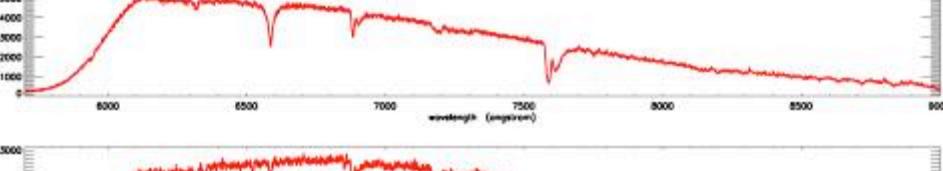
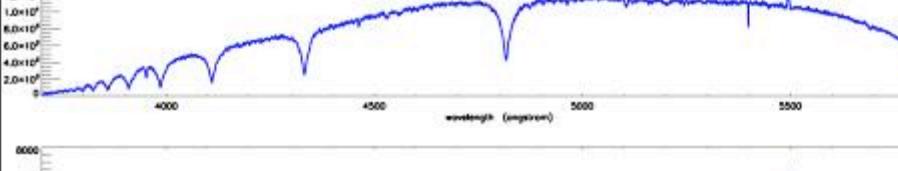
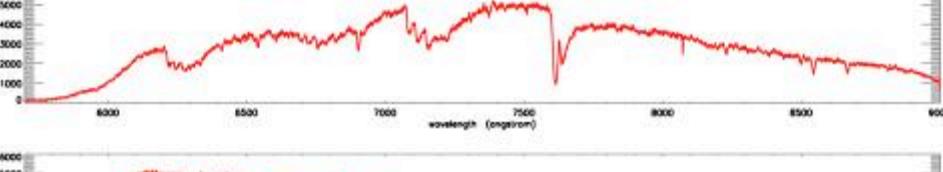
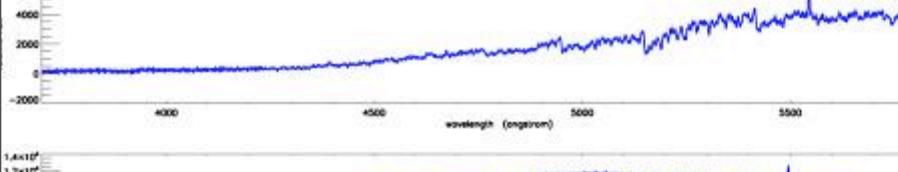
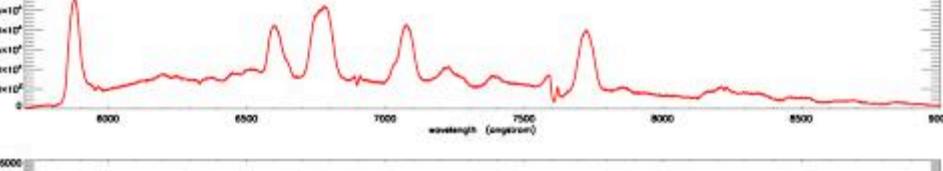
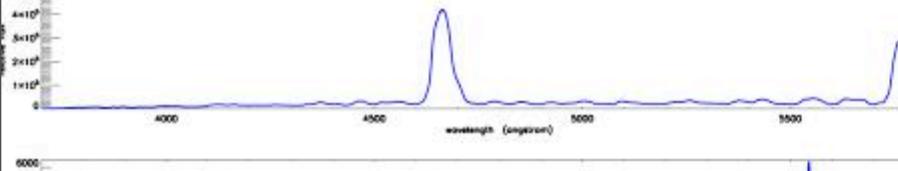
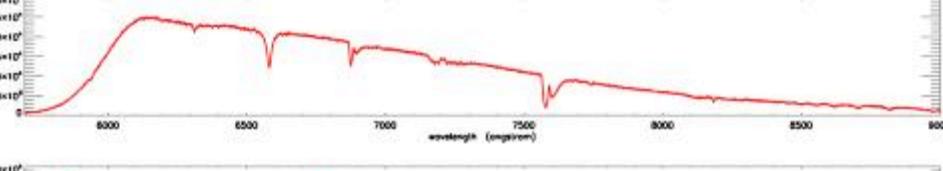
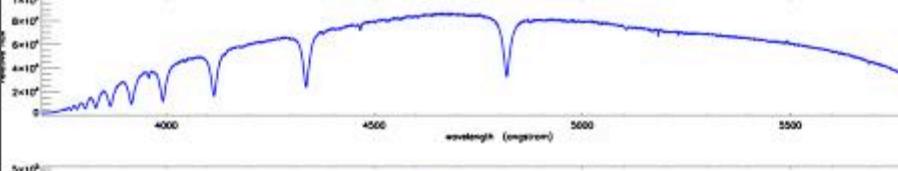
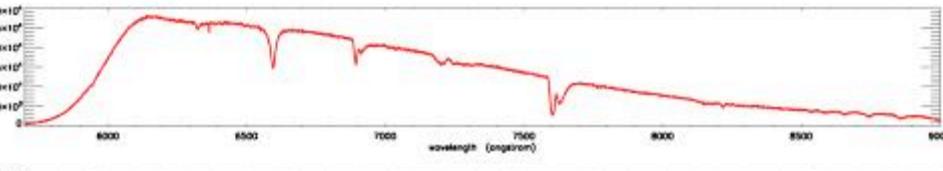
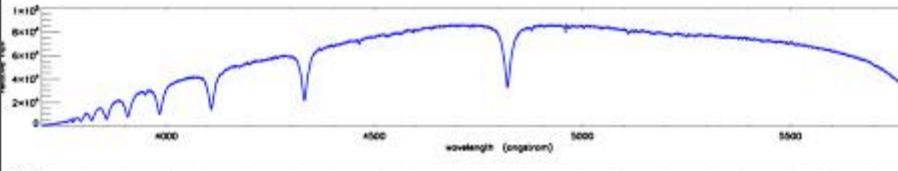
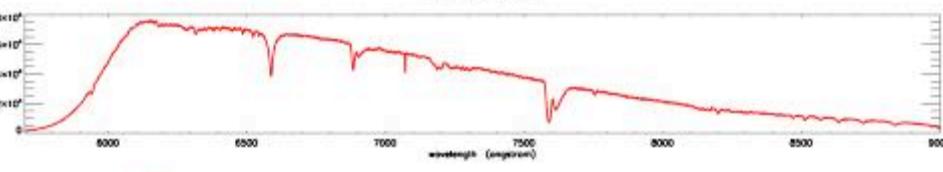
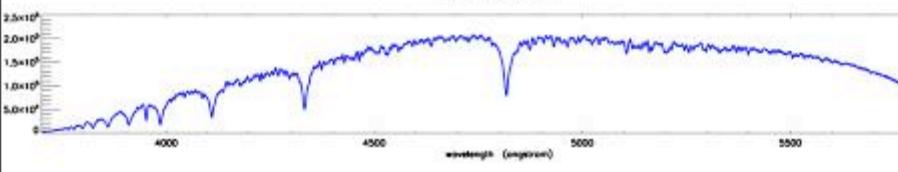
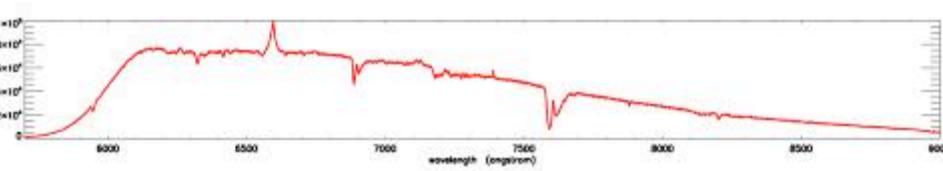
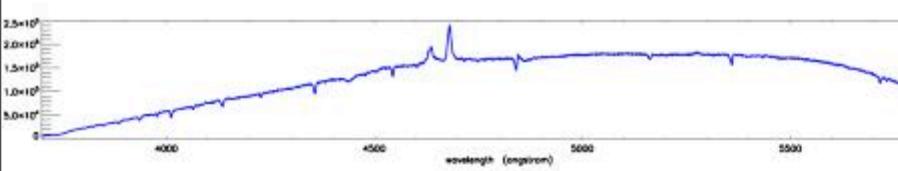
Spectra of stars (28/9/2008)

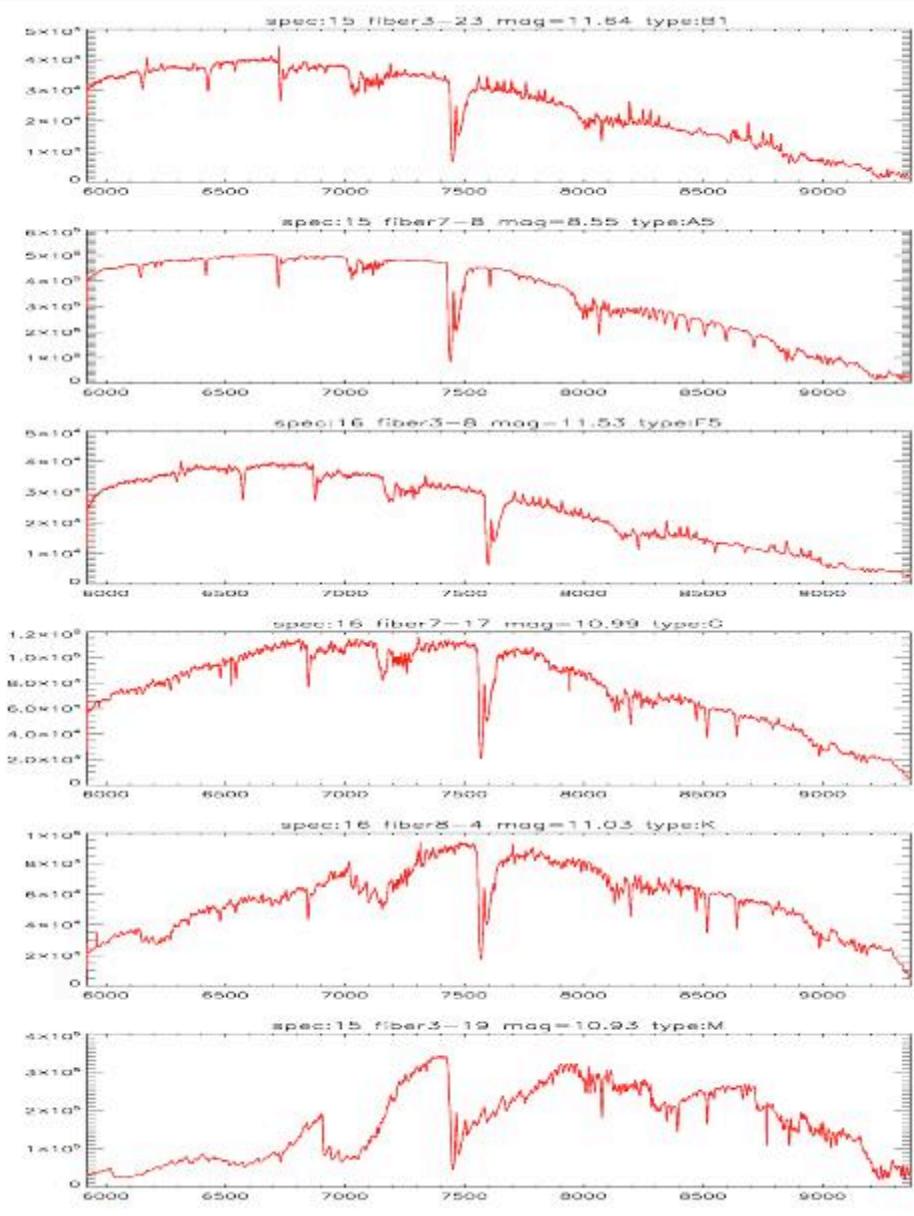
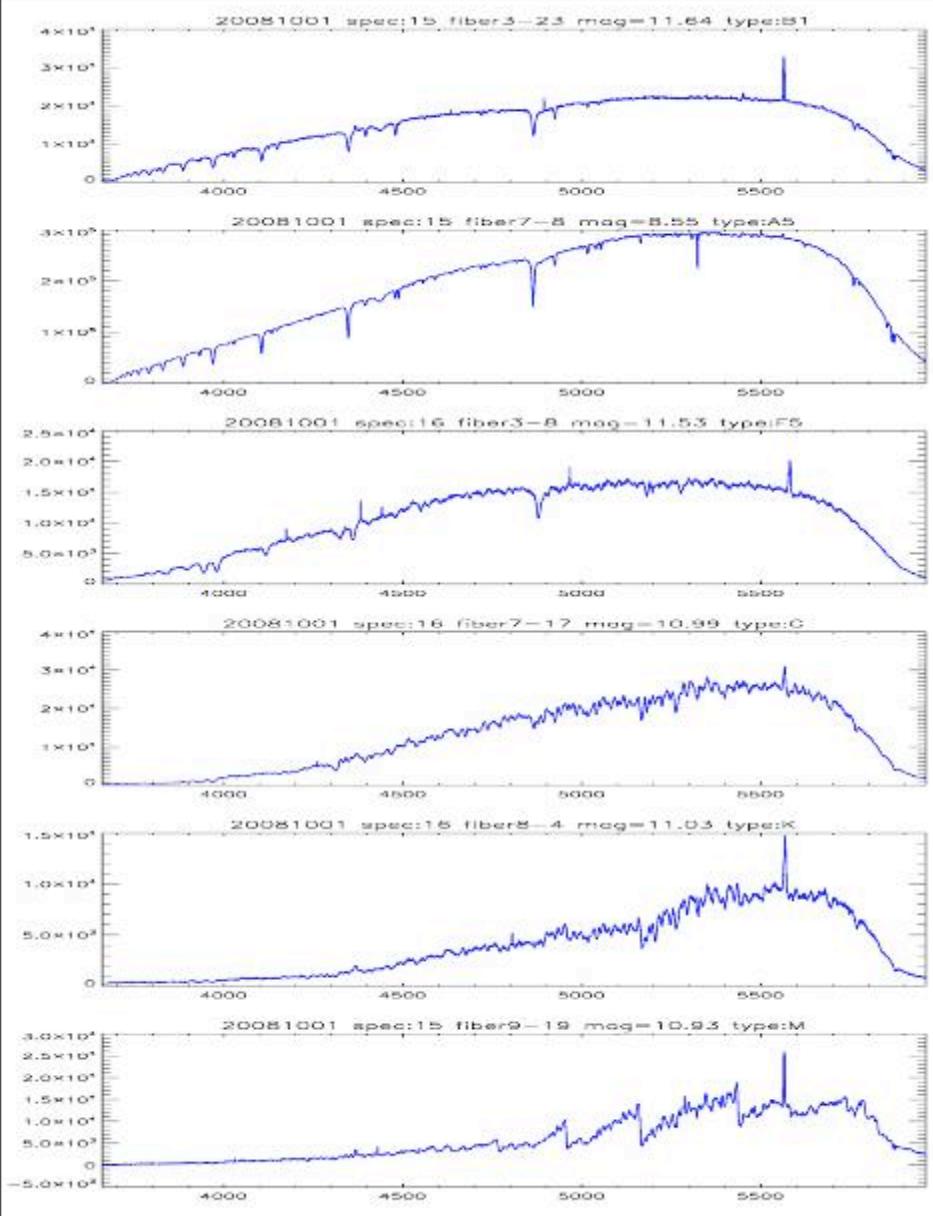


Red



Blue

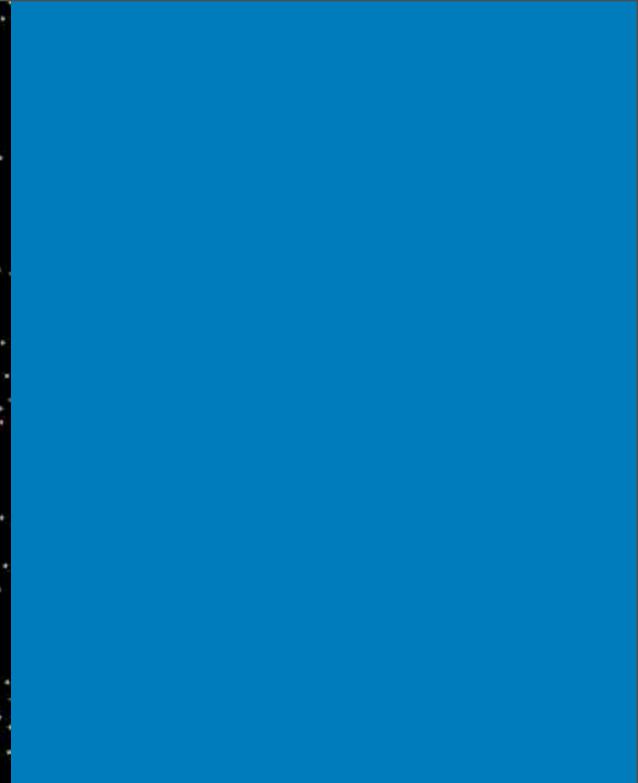
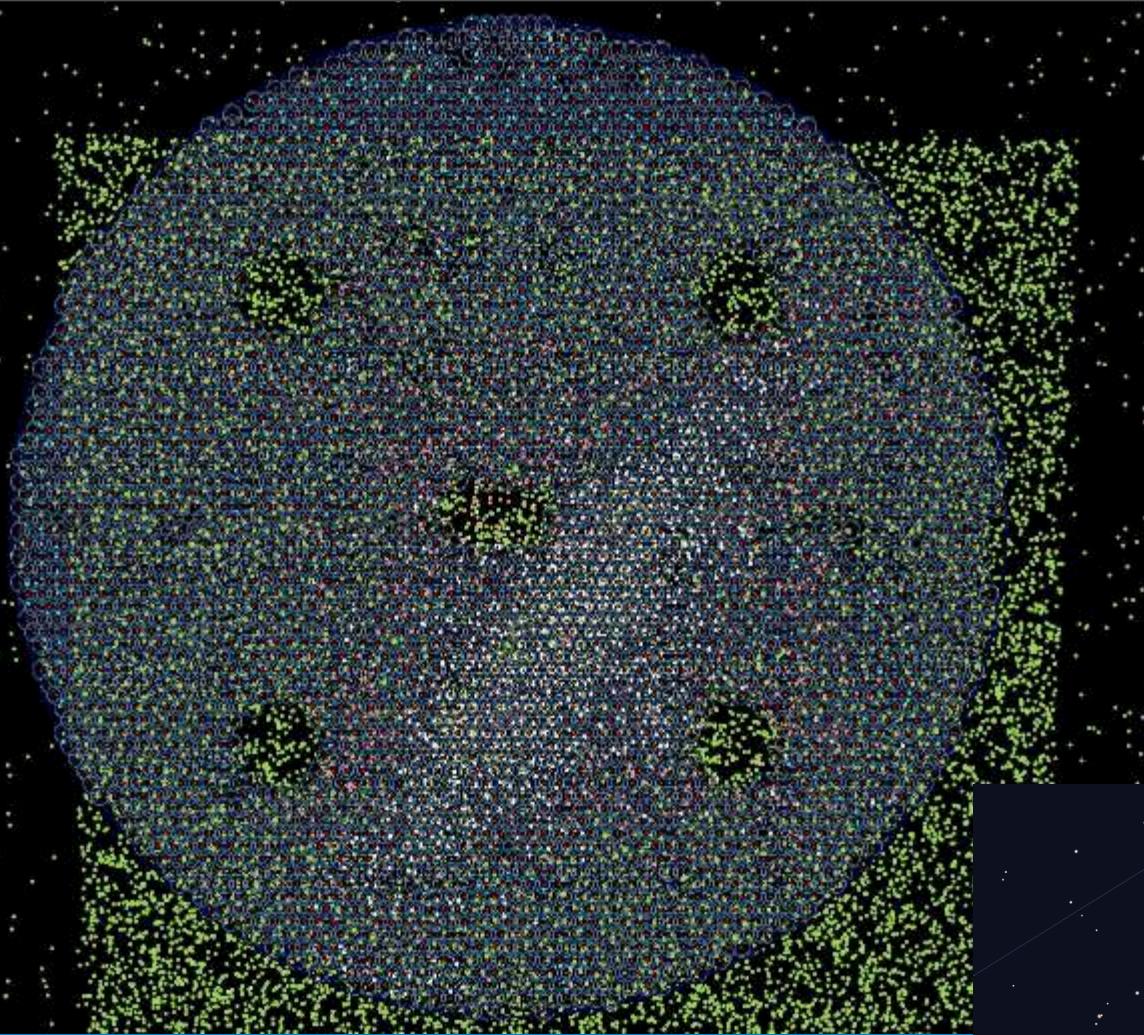


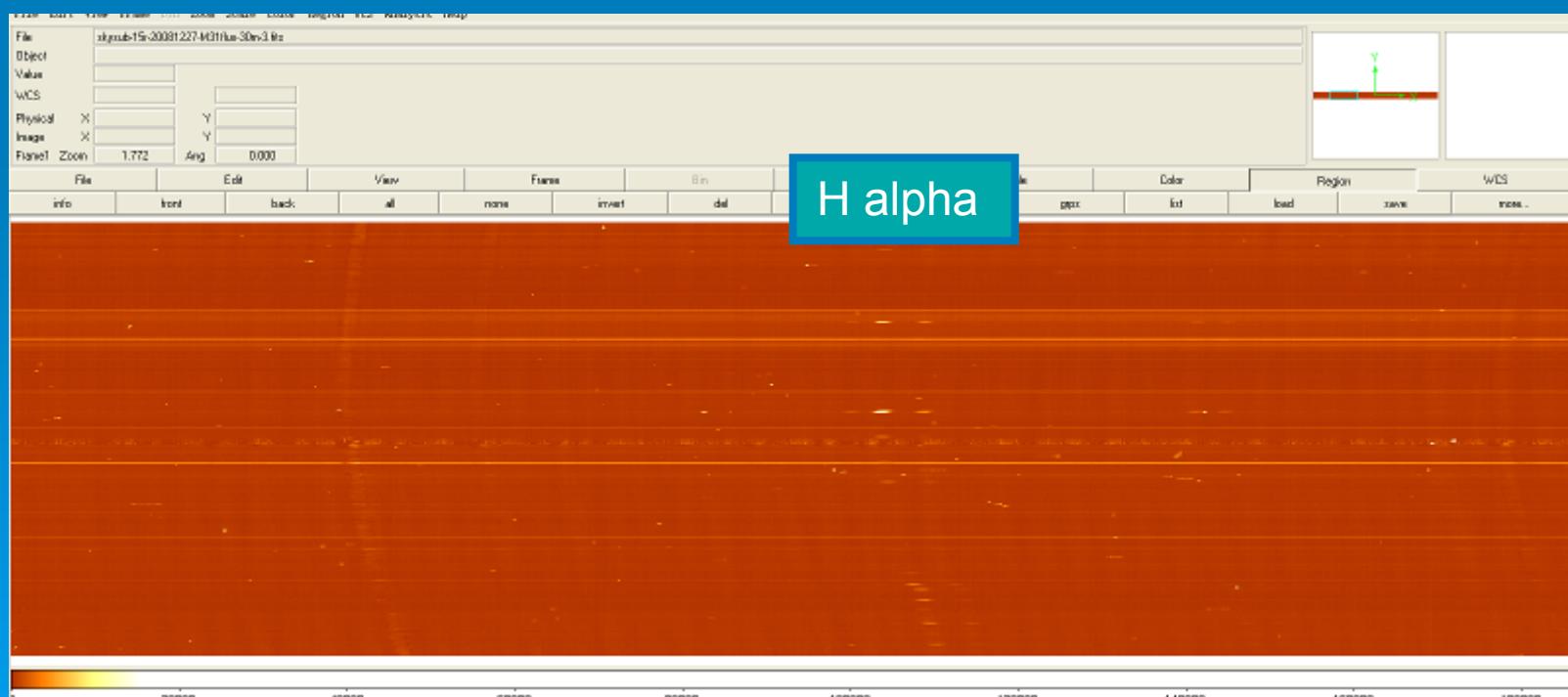
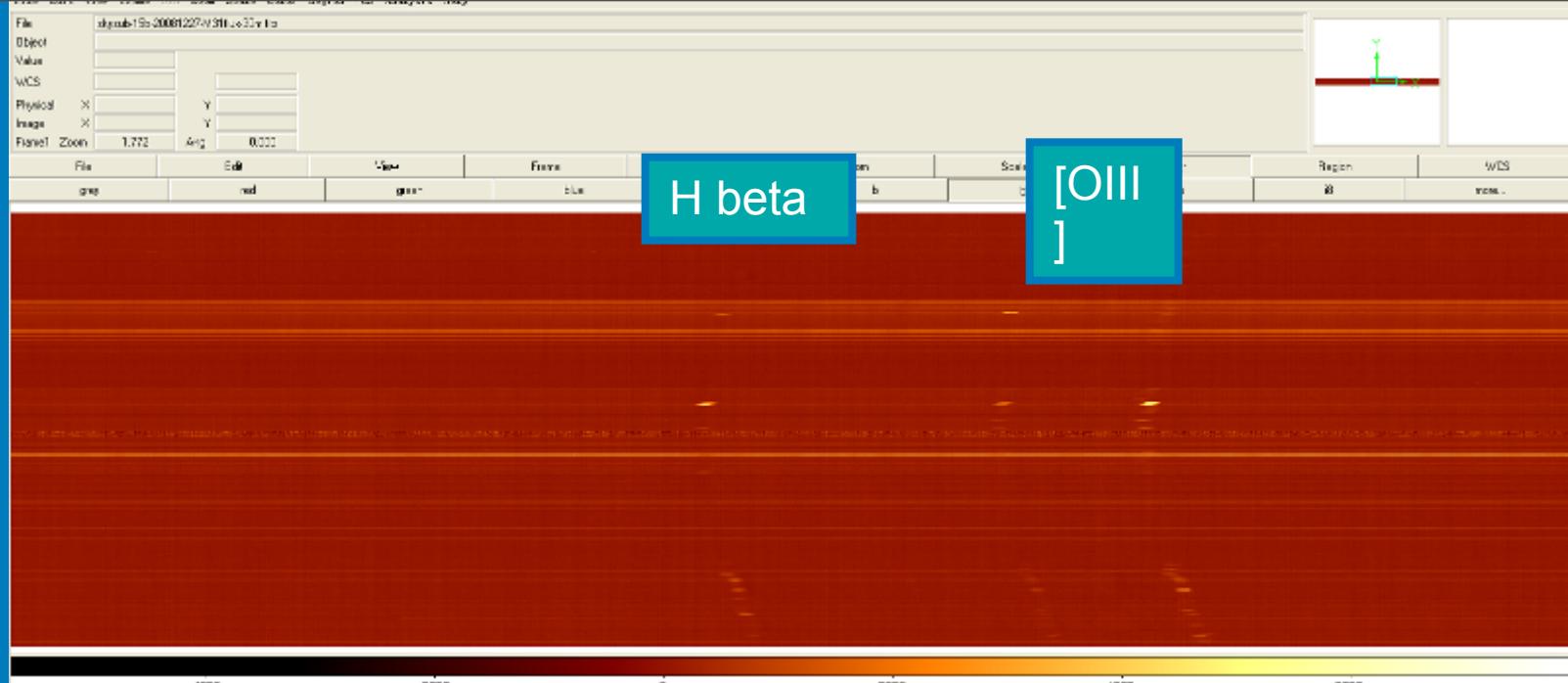


Test Observations

- Dec. 27, 2008
 - M31
 - Planetary nebula
 - Globular clusters
 - Others
 - Galaxies
 - Stars
 - 1800s Exp.

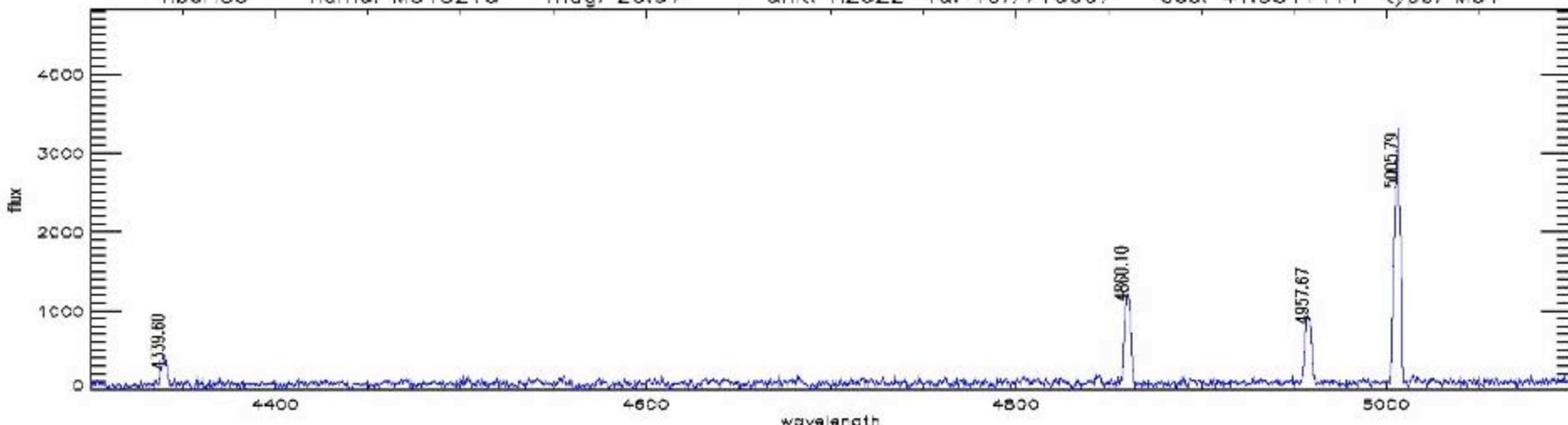




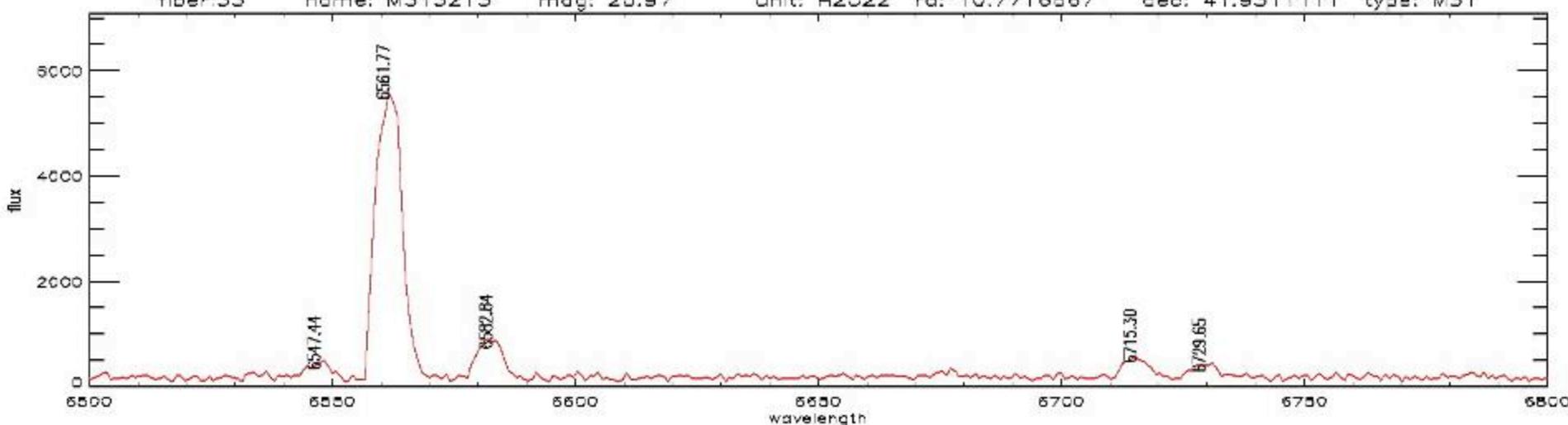


PN in M31

fiber:33 name: M313213 mag: 20.97 unit: H2022 ra: 10.7716867 dec: 41.9311111 type: M31



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Plan



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- 2009: commission period
 - Early Science

Plan

- 2009: commission period
 - Early Science
- 2010 / 2011: regular spectroscopic survey
 - 5 year survey

Spectroscopic Surveys

- Key projects include
 - extra-Galactic
 - Milky Way
 - cross-identification
- WG for the Milky Way study
- WG for extragalactic survey
 - Survey plan will be fixed in 2009

Galactic survey plan:

- Consider a 5 year survey plan
- spectra of 2.5m stars are expected
 - Using about $\frac{1}{2}$ of the total dark observing time for halo (down to 20m)
 - Using about $\frac{1}{2}$ of the grey and all bright nights for bright stars (18m) in the ‘green’ fields

Galactic Astronomy

- Dwarf galaxies and stellar moving groups;
- Constraining the gravitational potential;
- Search for extremely metal poor stars;
- The structure of the thin/thick disks;
- Survey of Galactic globular and open clusters.

Extragalactic Survey

Observe galaxies and QSOs
in the NGC of 7700 deg² (SDSS sky)
in the SGC of about 3500 deg²

- Shallow Survey:
2.3 million galaxies of $r < 18.8$ in 7500 deg²;
- Deep Galaxy Survey:
3.2 million galaxies of $r < 19.5$ in 3700 deg²
- Early Massive Galaxy Survey (EMG):
1 million galaxies of $V < 20.0$;
- QSO survey
0.6 million QSO of $i < 20.5$;

A data base of 7 million extragalactic spectra

Thank You !

